Machine Assisted Thought

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Collaborators

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Lee Dirks

Three Papers

- 2012arXiv1209.1318K
 - Finding and Recommending Scholarly Articles
- 2011ApSSP...1...23K
 - The Emerging Scholarly Brain
- 1993ASSL..182...21K
 - Advice from the Oracle: Really Intelligent Information Retrieval

Numbers

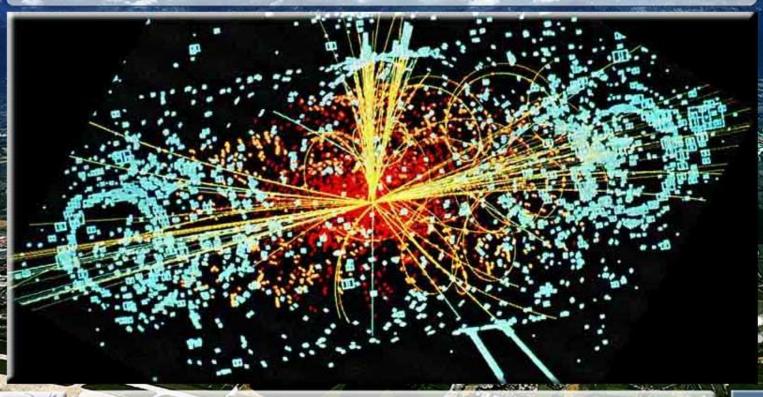
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LHC: The Large Hadron Collider

CMS -- Le "Compact Muon Solenoid"

Cette image montre une collision simulée d'une collision du CMS. Le centre de l'image montre où les protons sont entres en collision et l'énergie résultante d'annihilation produit des jets de nouvelles particules qui peuvent se déplacer dans le détecteur.

L'image est une de celles que nous espérons voir quand CMS sera en fonction: elle met en évidence le boson de Higgs, la particule qui confère une masse a toutes les autres particules et que le LHC devrait pouvoir détecteur.



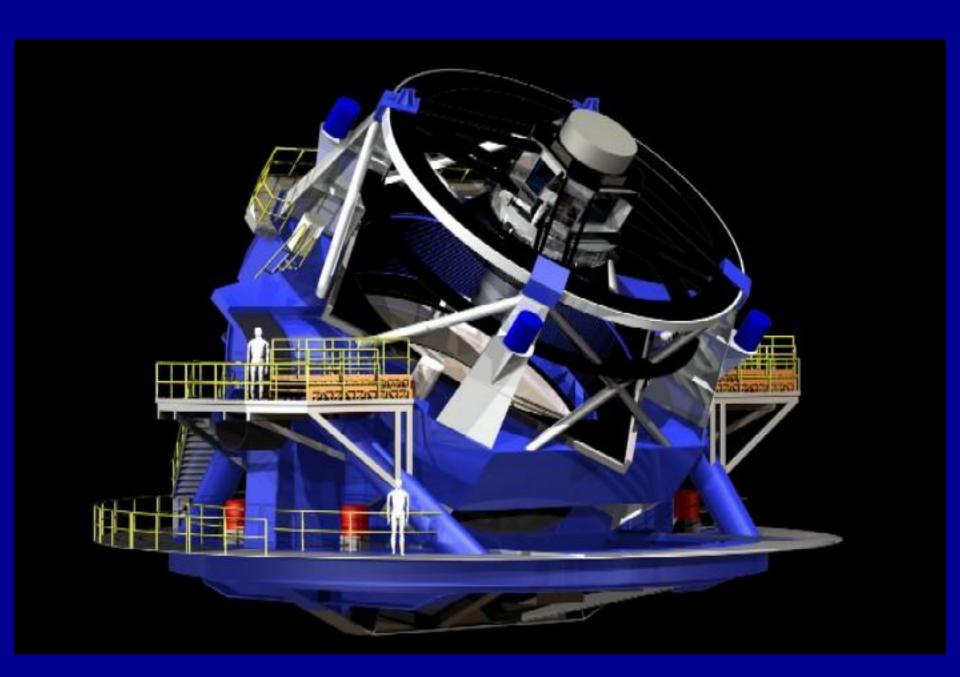
SPS LHC Ring Ring

CMS

Détecteur CMS: Récupération des données

1: CMS, 2: ATLAS, 3: LHCb, 4: ALICE





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Immanuel Kant,

Professor in Ronigsberg,

der Königt Academie der Wiffenschaften in Berlin

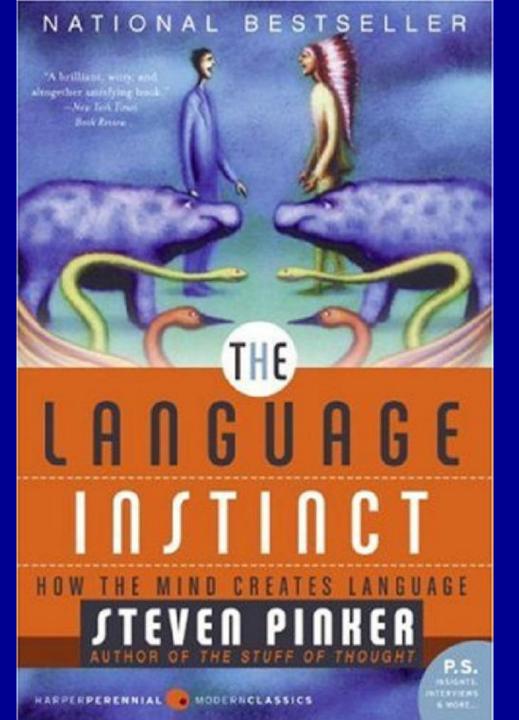


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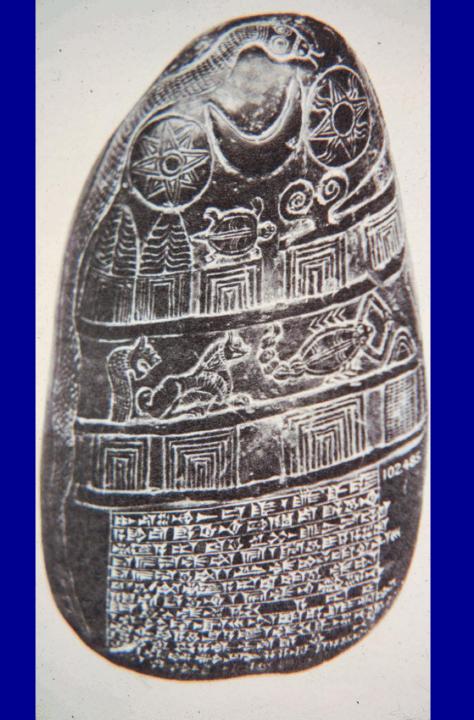
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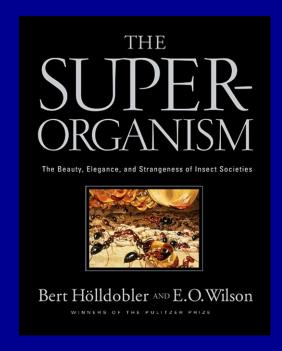
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Super-organisms



... the highest level of the ant colony is the totality of its membership rather than a particular set of superordinate individuals who direct the activity of members at lower levels.

Höldobler and Wilson (1990)





These are termites







WORKER

Actual size 1/4-inch

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Actual size 5/16-inch

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Actual size 1/2-inch

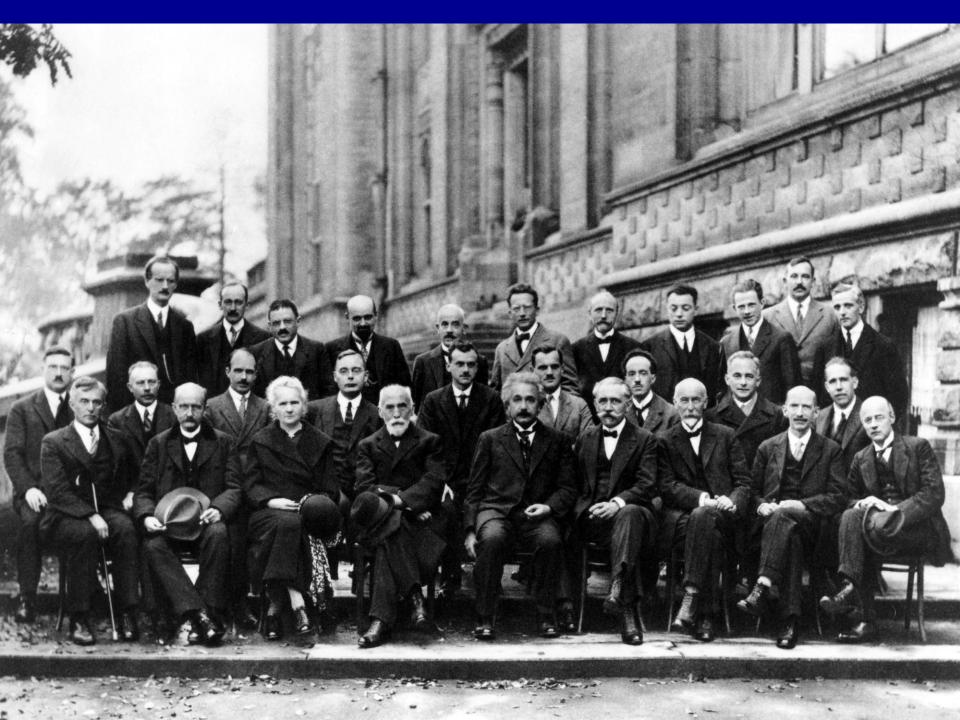


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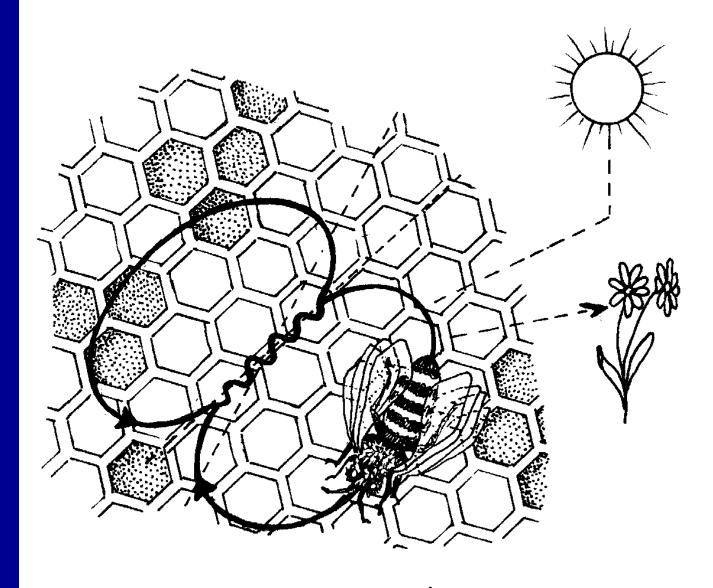


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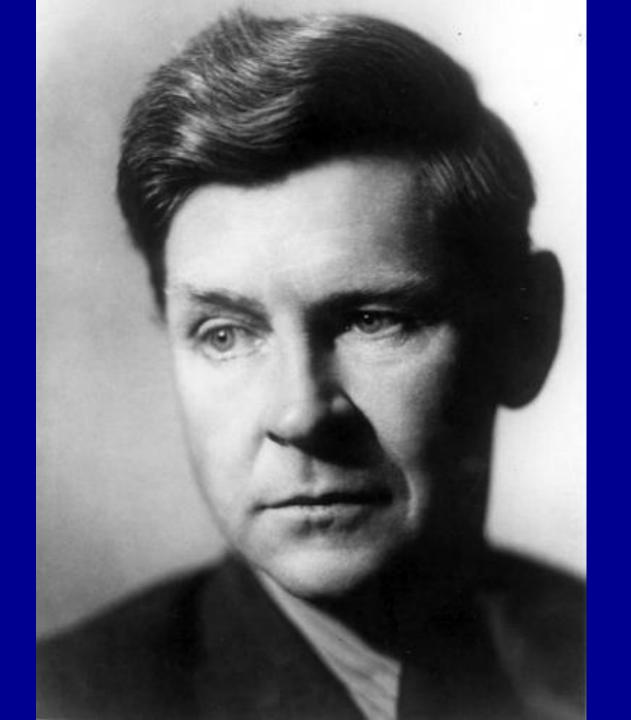


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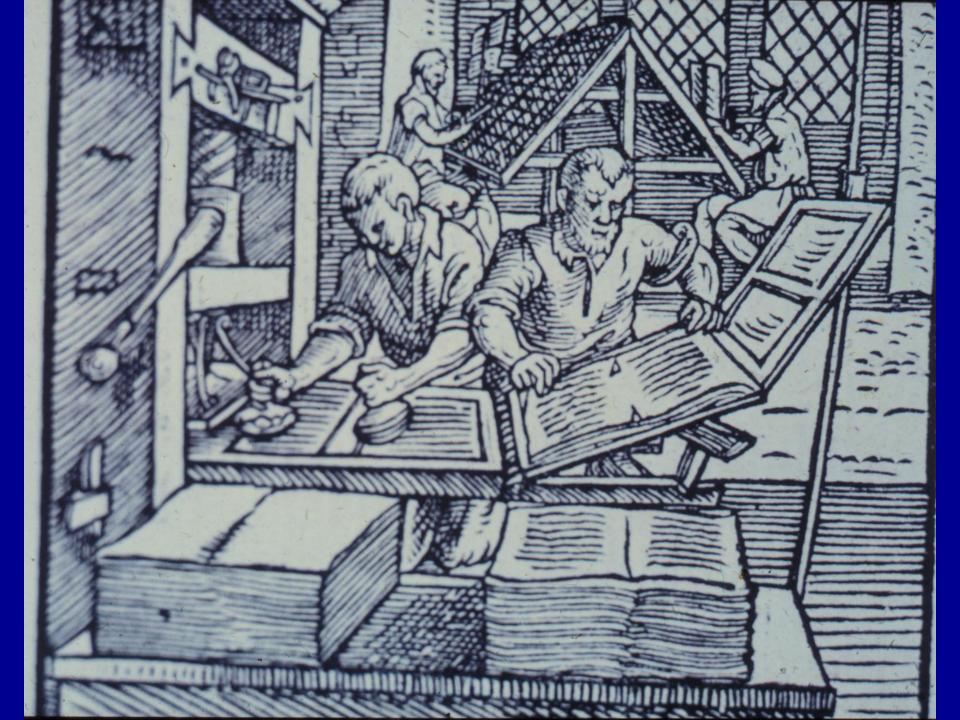


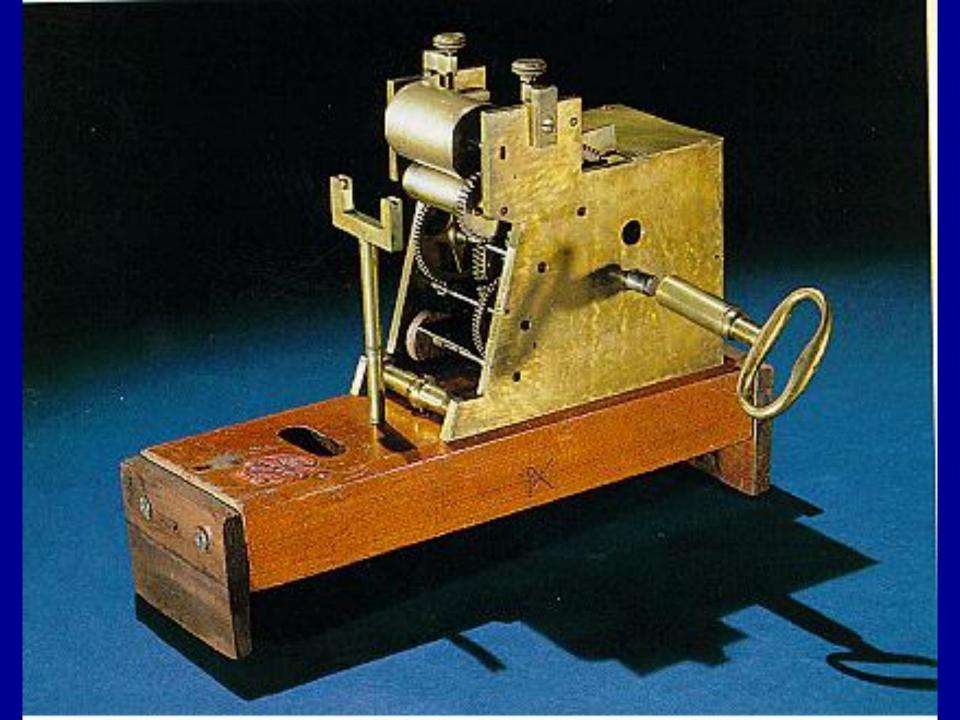


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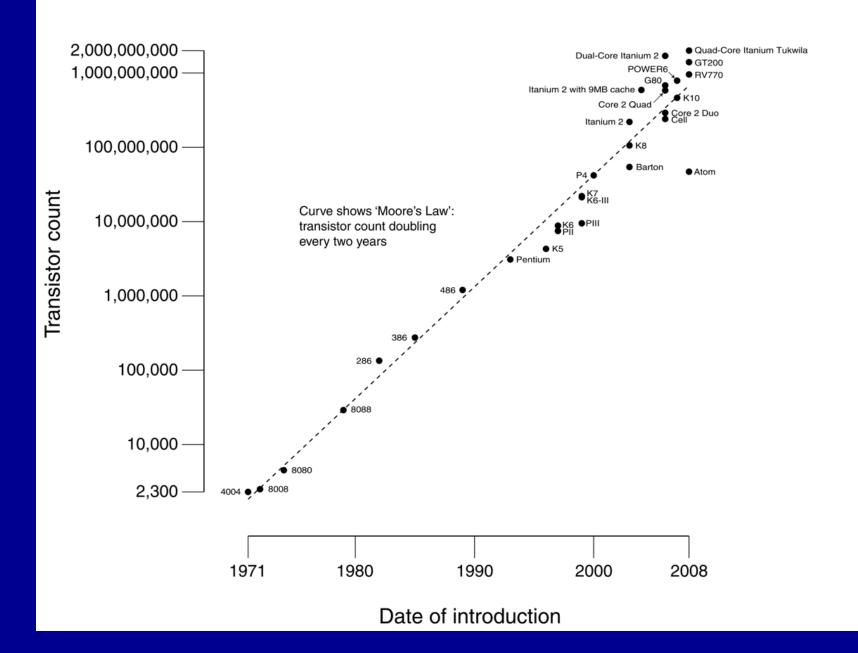


- The system of radiation which embraces the whole planet, and includes the million million brains of the race, becomes the physical basis of a racial self...
- But chiefly the racial mind transcends the minds of groups and individuals in philosophical insight into the true nature of space and time, mind and its objects, cosmical striving and cosmical perfection....
- For all the daily business of life, then, each of us is mentally a distinct individual, though his ordinary means of communication with others is "telepathic." But frequently he wakes up to be a groupmind...
- Of this obviously, I can tell you nothing, save that it differs from the lowlier state more radically than the infant mind differs from the mind of the individual adult, and that it consists of insight into many unsuspected and previously inconceivable features of the familiar world of men and things.





CPU Transistor Counts 1971-2008 & Moore's Law









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The New York Times

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"It's a very modest seeing and the our giving to sail you it's eat," the officer total "It will be a contact burtle, and it will be by

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SENATE LEADERS SEE PATH TO AVERT MANDATORY CUTS

FACING DAX, I DEADLINE

Bipartisan Support for a 3-Step Plan to Arreid the Fiscal Cliff

By SOMETHING WELLKOOM

WASHINGTON - Brook Mars of the Atlanta was really be dealing with the Theat of The facing the steadyr in Assess opting in the far see a problem to motion of Congress to Sweet spreament on a comprehensive tarked makes from their carbon from a liferal trigging to Europein.

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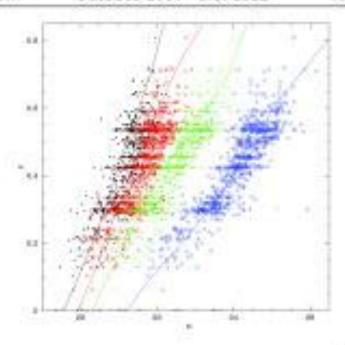
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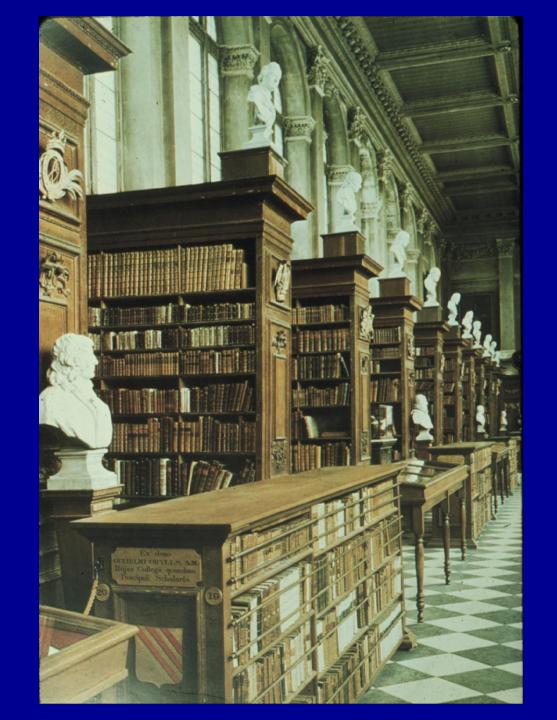
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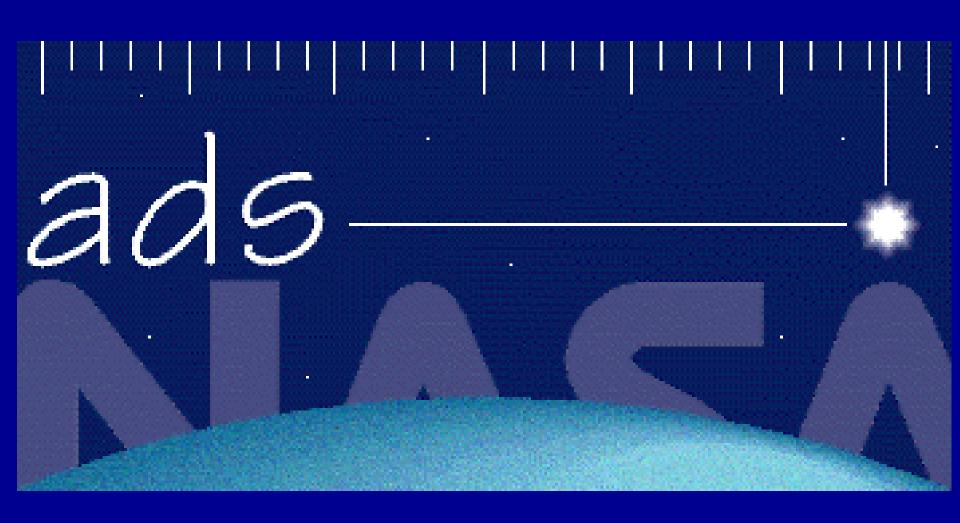


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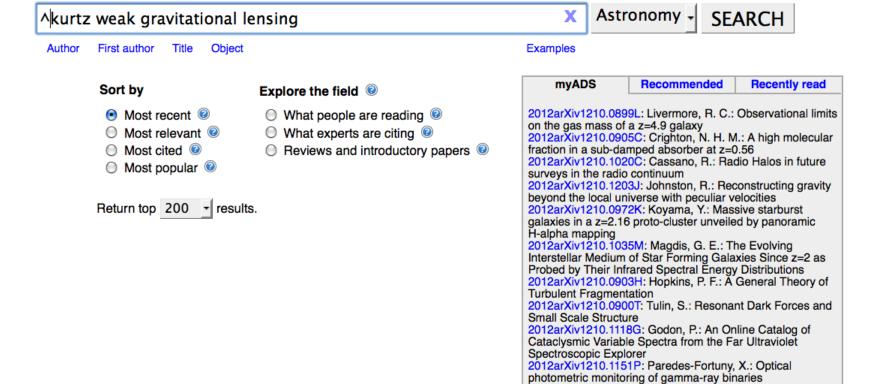
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2007ApJ...668..643L Limousin,+: Combining Strong and Weak Gravitational Lensing in Abell

2010ApJ...715..743K Kelly,+: Hubble Residuals of Nearby Type Ia Supernovae are Correlated with Host Galaxy Masses

2011arXiv1110.1377K Kelly,+: Core-Collapse Supernovae and Host Galaxy Stellar **Populations**

2011A&ARv..19...47K Kneib,+: Cluster lenses 1997ApJ...490..493N Navarro,+: A Universal Density Profile from Hierarchical Clustering









Testing Weak-lensing Maps with Redshift Surveys: A Subaru Field

Kurtz, Michael J.; Geller, Margaret J.; Utsumi, Yousuke; Miyazaki, Satoshi; Dell'Antonio, Ian P.; Fabricant, Daniel G.

show affiliations

The Astrophysical Journal, Volume 750, Issue 2, article id. 168 (2012). Published in May 2012

DOI: 10.1088/0004-637X/750/2/168

We use a dense redshift survey in the foreground of the Subaru GTO2deg2 weak-lensing field (centered at $\alpha 2000 = 16h04m44s$; $\delta 2000 = 43^{\circ}11'24''$) to assess the completeness and comment on the purity of massive halo identification in the weak-lensing map. The redshift survey (published here) includes 4541 galaxies: 4405 are new redshifts measured with the Hectospec on the MMT. Among the weak-lensing peaks with a signal-to-noise greater than 4.25, 2/3 correspond to individual massive systems; this result is essentially identical to the Geller et al. test of the Deep Lens Survey (DLS) field F2. The Subaru map, based on images in substantially better seeing than the DLS, enables detection of less massive halos at fixed redshift as expected. We demonstrate that the procedure adopted by Miyazaki et al. for removing some contaminated peaks from the weak-lensing map improves agreement between the lensing map and the redshift survey in the identification of candidate massive systems.

Kevwords:

Astronomy: cosmology: observations, galaxies: clusters: general, galaxies: distances and redshifts, gravitational lensing: weak, large-scale structure of Universe

arXiv: arXiv:1102.5743



Hectospec, the MMT's 300 Optical Fiber-Fed Spectrograph



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Abstract

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A Spectroscopic Catalog of 10 Distant Rich Clusters of Galaxies

Dressler, Alan; Smail, Ian; Poggianti, Bianca M.; Butcher, Harvey; Couch, Warrick J.; Ellis, Richard S.; Oemler, Augustus, Jr.

The Astrophysical Journal Supplement Series, Volume 122, Issue 1, pp. 51-80. Published in May 1999

DOI: 10.1086/313213

We present spectroscopic observations of galaxies in the fields of 10 distant clusters for which we have previously presented deep imaging with WFPC2 on board the Hubble Space Telescope. The clusters span the redshift range z=0.37-0.56 and are the subject of a detailed ground- and space-based study to investigate the evolution of galaxies as a function of environment and epoch. The data presented here include positions, photometry, redshifts, spectral line strengths, and classifications for 657 galaxies in the fields of the 10 clusters. The catalog is composed of 424 cluster members across the 10 clusters and 233 field galaxies, with detailed morphological information from our WFPC2 images for 204 of the cluster galaxies and 71 in the field. We illustrate some basic properties of the catalog, including correlations between the morphological and spectral properties of our large sample of cluster galaxies. A direct comparison of the spectral properties of the high-redshift cluster and field populations suggests that the phenomenon of strong Balmer lines in otherwise passive galaxies (commonly called E+A but renamed here as the k+a class) shows an order-of-magnitude increase in the rich cluster environment compared with a more modest increase in the field population. This suggests that the process or processes involved in producing k+a galaxies are either substantially more effective in the cluster environment or that this environment prolongs the visibility of this phase. A more detailed analysis and modeling of these data is presented in Poggianti et al.

Actronomy: GALAYIES: CLUSTEDS: GENERAL GALAYIES: DISTANCES AND DERSHIETS

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Data Products

SIMBAD objects (675) NED objects (674) Archival data (3)

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2000ApJ...534L.131J Joffre,+: Weak Gravitational Lensing by the Nearby Cluster Abell 3667

2011ApJ...733L...30H Hildebrandt,+: Lensing Magnification: A Novel Method to Weigh High-redshift Clusters and its Application to SpARCS

2009ApJ...705..809T Tran,+: A Spectroscopically Confirmed Excess of 24 μ m Sources in a Super Galaxy Group at z = 0.37: Enhanced Dusty Star Formation Relative to the Cluster and Field Environment

2010ApJ...715..743K Kelly,+: Hubble Residuals of Nearby Type Ia Supernovae are Correlated with Host Galaxy Masses

2011arXiv1110.1377K Kelly,+: Core-Collapse Supernovae and Host Galaxy Stellar Populations

2011ARA&A..49..409A Allen,+: Cosmological Parameters from Observations of Galaxy Clusters

1996A&AS..117..393B Bertin,+: SExtractor: Software for source extraction.

1998ApJ...500..525S Schlegel,+: Maps of Dust Infrared Emission for Use in Estimation of Reddening and Cosmic Microwave Background Radiation Foregrounds

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A SPECTROSCOPIC CATALOG OF 10 DISTANT RICH CLUSTERS OF GALAXIES

ALAN DRESSLER, IAN SMAIL, 2,3 BIANCA M. POGGIANTI, 4,5,6 HARVEY BUTCHER, WARRICK J. COUCH, RICHARD S. ELLIS, AND AUGUSTUS OEMLER, JR. Received 1998 June 12; accepted 1998 December 21

ABSTRACT

We present spectroscopic observations of galaxies in the fields of 10 distant clusters for which we have previously presented deep imaging with WFPC2 on board the Hubble Space Telescope. The clusters span the redshift range z = 0.37 - 0.56 and are the subject of a detailed ground- and space-based study to investigate the evolution of galaxies as a function of environment and epoch. The data presented here include positions, photometry, redshifts, spectral line strengths, and classifications for 657 galaxies in the fields of the 10 clusters. The catalog is composed of 424 cluster members across the 10 clusters and 233 field galaxies, with detailed morphological information from our WFPC2 images for 204 of the cluster galaxies and 71 in the field. We illustrate some basic properties of the catalog, including correlations between the morphological and spectral properties of our large sample of cluster galaxies. A direct comparison of the spectral properties of the high-redshift cluster and field populations suggests that the phenomenon of strong Balmer lines in otherwise passive galaxies (commonly called E + A but renamed here as the k + a class) shows an order-of-magnitude increase in the rich cluster environment compared with a more modest increase in the field population. This suggests that the process or processes involved in producing k + a galaxies are either substantially more effective in the cluster environment or that this environment prolongs the visibility of this phase. A more detailed analysis and modeling of these data is presented in Poggianti et al.

Subject headings: galaxies: clusters: general — galaxies: distances and redshifts — galaxies: evolution — galaxies: photometry

1. INTRODUCTION

The change with redshift observed in the proportion of star-forming galaxies in the cores of rich clusters was uncovered over 20 years ago, by Butcher & Oemler (1978, 1984), but it remains one of the clearest and most striking examples of galaxy evolution. Considerable effort has gone into acquiring photometric information that would elucidate the physical processes active in distant clusters and their effects on the evolution of both the star-forming (Lavery & Henry 1994; Lubin 1996; Rakos & Schombert 1995; Rakos, Odell, & Schombert 1997) and passive galaxies (Aragón-Salamança et al. 1993; Stanford, Eisenhardt, & Dickinson 1995, 1998; Smail et al. 1998). Further impetus has been provided by observations of the recent transformation of the S0 population of clusters (Dressler et al. 1997), which may allow a closer connection to be drawn between the galaxy populations of distant clusters and the evolutionary signatures found in their local universe counterparts (Caldwell & Rose 1997; Bothun & Gregg 1990).

However, it was the advent of spectroscopic surveys of the distant cluster populations (e.g., Dressler & Gunn 1983, 1992, hereafter DG92; Couch & Sharples 1987, hereafter

CS87: Barger et al. 1996: Abraham et al. 1996: Fisher et al. 1998) that uncovered the real breadth of the changes in galaxies in these environments, including several spectral signatures of evolutionary change, such as evidence for a strong decline in the star formation rates of many cluster galaxies in the recent past. The advent of high-spatial resolution imaging with the Hubble Space Telescope (HST) provided a further breakthrough, giving morphological information on the galaxies in these distant clusters. This could be used to link the evolution of stellar populations in the galaxies with the evolution of their structure in order to understand how the various galaxy types we see in the local universe came to be. Pre- and postrefurbishment HST observations by two groups (Couch et al. 1994, 1998; Dressler et al. 1994; Oemler, Dressler, & Butcher 1997) were used in early attempts to correlate spectral evolution with morphological/structural data and to provide some insight into the mechanisms that might be driving the strong evolution in the cluster galaxy population. These two programs were extended from cycle 4 into the "MORPHS" project, which accumulated postrefurbishment WFPC2 images for 11 fields in 10 clusters at z = 0.37-0.56, viewed at a time some 2-4 h⁻¹ billion yr before the present day. The photometric and morphological galaxy catalogs from these images were presented in Smail et al. (1997a, hereafter S97), while the data have also been used to study the evolution of the early-type galaxies within the clusters, using both color (Ellis et al. 1997) and structural information (Barger et al. 1998), the evolution of the morphology-density relation of

¹ The Observatories of the Carnegie Institution of Washington, 813 Santa Barbara Street, Pasadena, CA 91101-1292.

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⁵ Royal Greenwich Observatory, Madingley Road, Cambridge CB3 0EZ, UK.

⁶ Osservatorio Astronomico di Padova, vicolo dell'Osservatorio 5, 35122 Padova, Italy.

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⁸ School of Physics, University of New South Wales, Sydney 2052, Australia.

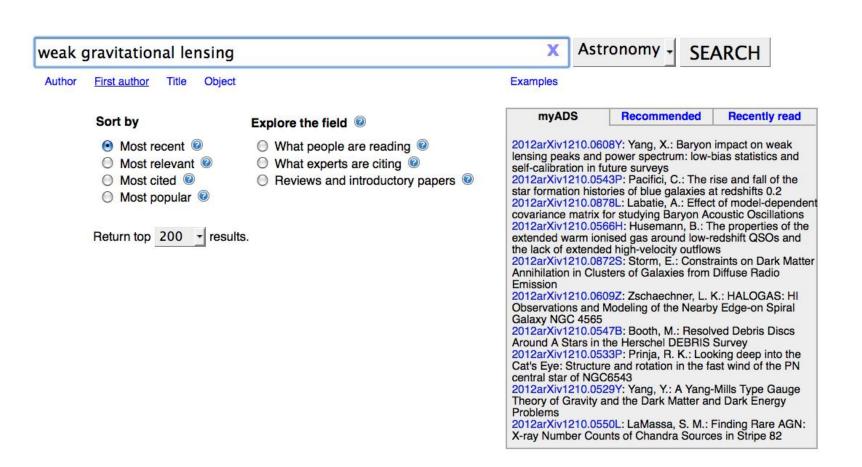
 $^{^9}$ We use $q_0=0.5$ and $h=H_0/100~{\rm km~s^{-1}~Mpc^{-1}}.$ For this geometry 1'' is equivalent to 3.09 h^{-1} kpc for our lowest redshift cluster and 3.76 h^{-1} kpc for the most distant,

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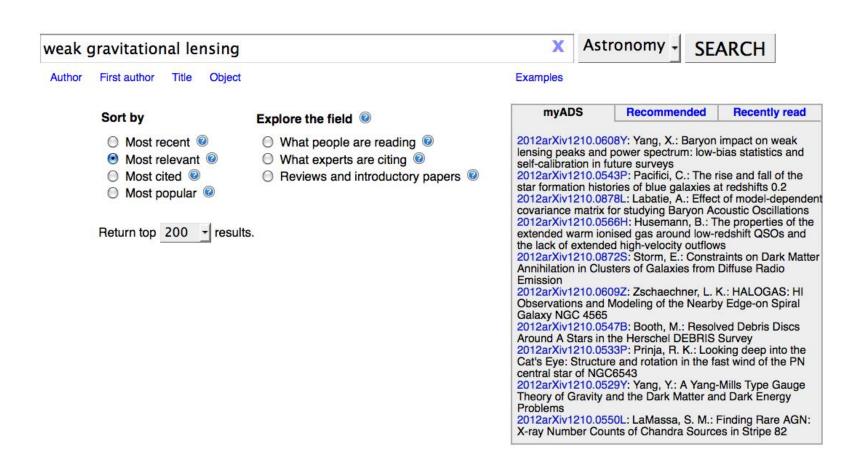








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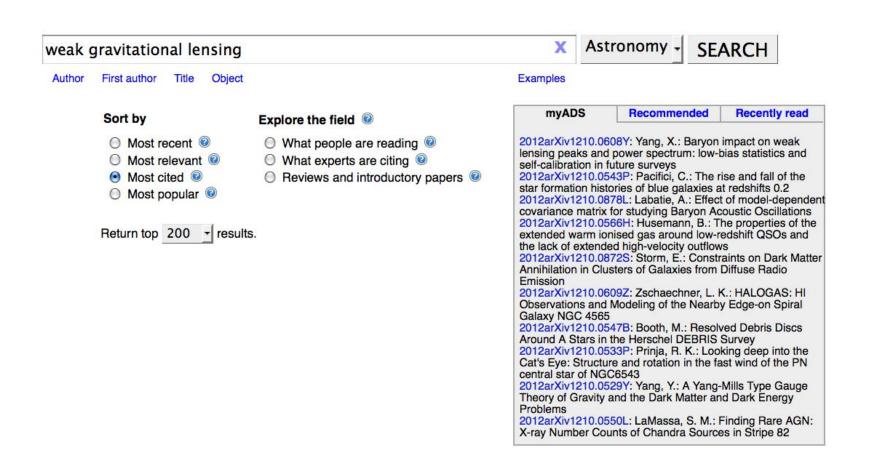








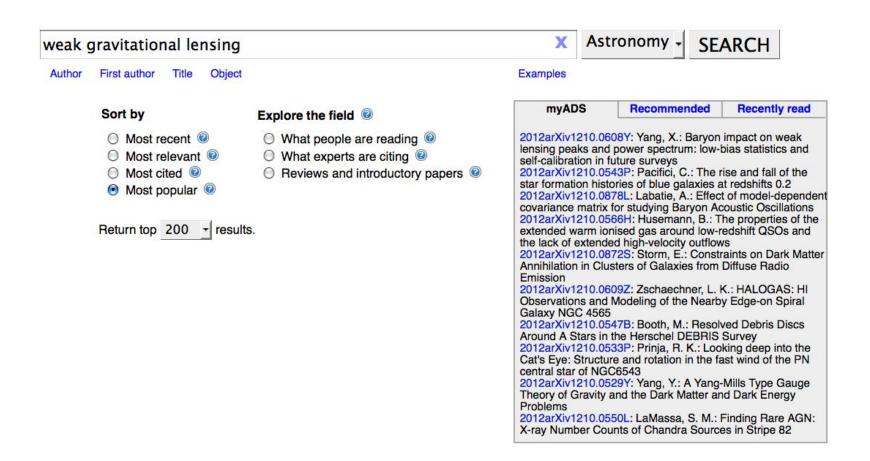
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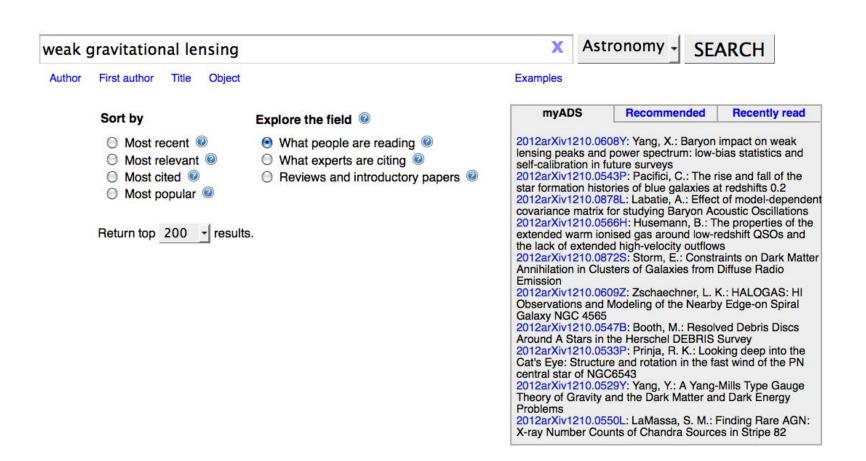
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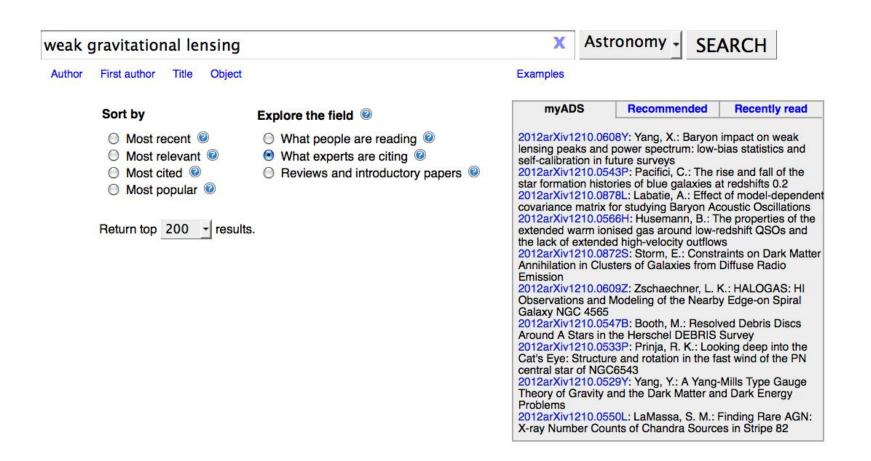








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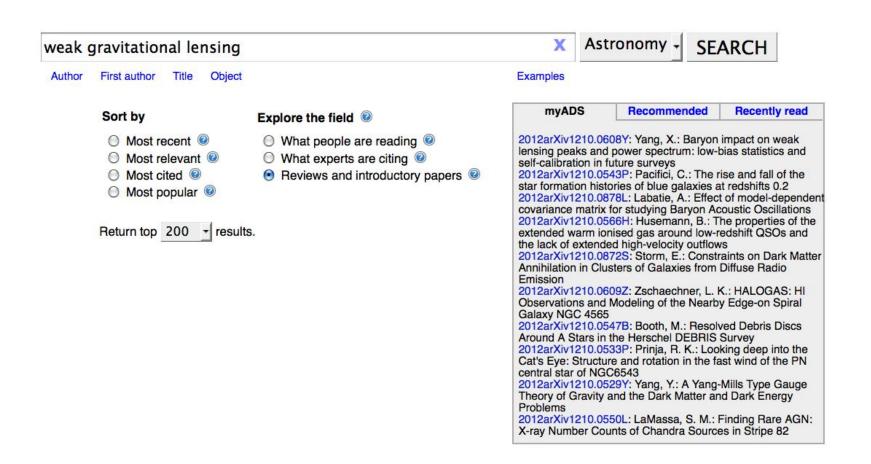








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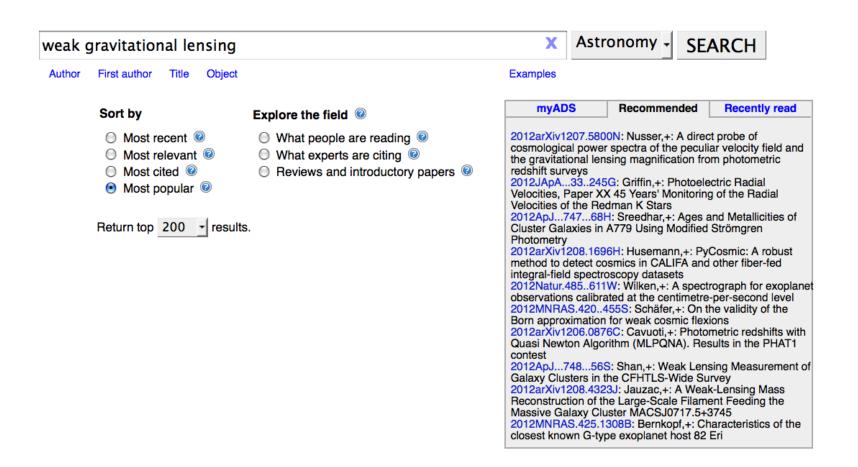








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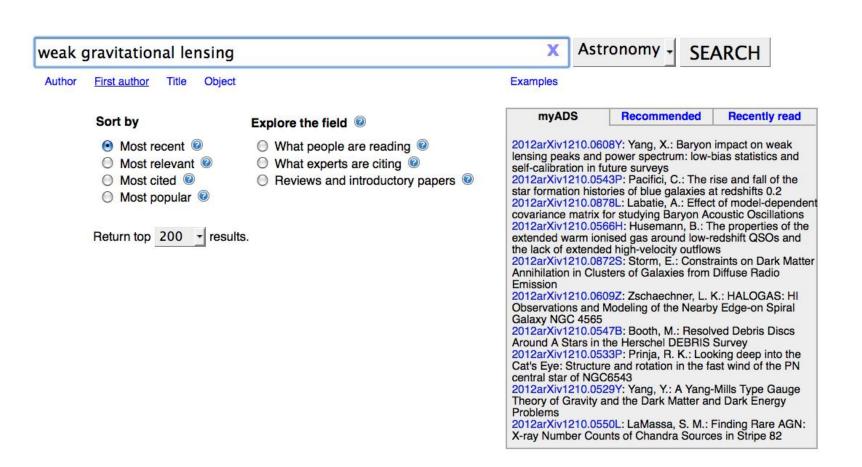


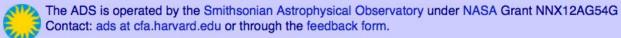
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The Astrophysical Journal, Volume 757, Issue 1, article id. 22 (2012).

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DOI: 10.1088/0004-637X/757/1/22

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lower limits on these values. The need for halo elongation can be partially obviated by non-thermal pressure support and, perhaps entirely, by systematic errors in the X-ray mass measurements. We estimate the effect of background structures based on MMT/Hectospec spectroscopic redshifts and find that these tend to lower M vir further by ~7% and

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- For all the daily business of life, then, each of us is mentally a distinct individual, though his ordinary means of communication with others is "telepathic." But frequently he wakes up to be a groupmind...
- Of this obviously, I can tell you nothing, save that it differs from the lowlier state more radically than the infant mind differs from the mind of the individual adult, and that it consists of insight into many unsuspected and previously inconceivable features of the familiar world of men and things.